THE EFFECT OF THE EARTH'S ROTATION ON THE VELOCITY OF LIGHT

PART I

By A. A. MICHELSON

ABSTRACT

Theory of the effect of the rotation of the earth on the velocity of light as derived on the hypothesis of a fixed ether.

Historical Remarks.—The theory was given originally in 1904. The experiment was undertaken at the urgent instance of Dr. L. Silberstein. A preliminary experiment at Mount Wilson in 1923 showed that it was necessary to resort to an exhausted pipeline.

In the Philosophical Magazine, (6) 8, 716, 1904, a plan was proposed for testing the effect of the earth's rotation on the velocity of light. The expression for the difference in path between two interfering pencils, one of which travels in a clockwise, and the other in a counterclockwise direction, may be deduced on the hypothesis of a fixed ether as follows:

If \( l_1 \) is the length of path at latitude \( \phi_1 \) and \( l_2 \) that at latitude \( \phi_2 \), \( v_1 \) and \( v_2 \) the corresponding linear velocities of the earth's rotation, and \( V \) the velocity of light, the difference in time required for the two pencils to return to the starting-point will be

\[
T = \frac{2 l_2 v_2}{V^2 - v_2^2} - \frac{2 l_1 v_1}{V^2 - v_1^2}
\]
or with sufficient approximation

\[ T = 2 \left( \frac{l_2v_2 - l_1v_1}{V^2} \right) \]

or if

\[ l_1 = l_0 \cos \phi_1 \quad \text{and} \quad v_1 = v_0 \cos \phi_1 \]
\[ l_2 = l_0 \cos \phi_2 \quad \text{and} \quad v_2 = v_0 \cos \phi_2 \]

and

\[ \phi_1 - \phi_2 = \frac{\hbar}{R} \]

\((R=\text{earth’s radius}),\)

the resulting difference in phase of the two pencils will be

\[ \Delta = \frac{4l\hbar}{V\lambda} \omega \sin \phi,^1 \]

in which \(\omega\) is the earth’s angular velocity, and \(\lambda\) the effective wavelength of the light employed.

The experiment remained in abeyance for many years, until, at the urgent instance of Dr. L. Silberstein, the writer was convinced of the importance of the work, notwithstanding serious difficulties which were anticipated in the way of raising the necessary funds. The greatest expense would be in arranging a pipe line a mile long and a foot in diameter, such as would be required for the work. In the hope that this device might not be necessary, however, it was decided to attempt the experiment in the open air at Mount Wilson.

The work was undertaken at Mount Wilson during the summer of 1923, with a circuit over one mile in length. The interference fringes between the two pencils, one of which traversed the circuit clockwise, and the other counterclockwise, were observed most clearly during the half-hour before and after sunset. But even under the best conditions, the interference fringes were so unsteady that it was found impossible to make any reliable measurements.

A doubt had been raised concerning the possibility of referring

^1 In the original article, \(\Delta\) was erroneously given as half of this value. Dr. L. Silberstein (Journal of the Optical Society, 5, 291, 1921) deduced the expression equivalent to the above, which was confirmed by Dr. A. C. Lunn, ibid., 6, 112, 1922, except that the area inclosed is substituted for the product \(l\hbar\).
any expected displacement to a fiducial zero with which to compare the results. Such a fiducial zero was furnished by providing a double circuit, in one of which the area, on which the expected displacement depends, was much greater than in the other. The impossibility of obtaining accurate results in the open-air experiments on Mount Wilson showed that it was clearly necessary to resort to a pipe line about one mile long and one foot in diameter which could be exhausted of air.

Funds for this experiment, amounting to about $17,000, were furnished by the University of Chicago, with an additional contribution of $491.55 made through the efforts of Dr. L. Silberstein. With this support it was decided to perform the experiments at Clearing, Illinois; and in this work, which is to be discussed in the next article, Dr. Gale was invited to join.
THE EFFECT OF THE EARTH'S ROTATION ON THE VELOCITY OF LIGHT

PART II

BY A. A. MICHELSON AND HENRY G. GALE

Assisted by FRED PEARSON

ABSTRACT

Experimental Test of Theory.—Air was exhausted from a twelve-inch pipe line laid on the surface of the ground in the form of a rectangle 2010×1113 feet. Light from a carbon arc was divided at one corner by a thinly coated mirror into direct and reflected beams, which were reflected around the rectangle by mirrors at the corners. The two beams returning to the original mirror produced interference fringes. The beam traversing the rectangle in a counter-clockwise direction was retarded. The observed displacement of the fringes was found to be 0.230±.005, agreeing with the computed value 0.236±.002 within the limits of experimental error.

A rectangular tract of land at Clearing, Illinois, 2010 feet from east to west and 1113 feet from north to south, was carefully surveyed and staked by Dr. Kannenstine, and twelve-inch water pipes were laid straight and level around the entire circuit with a double line across one end.

The general plan of the arrangement is shown in Figure 1. Cast-iron boxes at the corners contained the mirrors. Figure 2 shows the details of one of the corner boxes. Delicate screw-and-lever systems, operated from outside the boxes through carefully fitted beveled joints, rendered it possible to adjust the mirrors readily about horizontal or vertical axes. The boxes were set in heavy concrete piers, and connected to the pipes by flexible joints of canvas and rubber. Similar joints were inserted in the pipe lines, about four hundred feet apart, and served as expansion joints.

The plane-parallel plates at A and B were lightly coated with gold, and that at C with silver, to reflect and transmit the desired proportions of light. The mirrors at D, E, and F were heavily silvered. A telephone system, consisting of portable sets kindly loaned by the Chicago Telephone Company, made it possible for an observer at A to direct assistants, one at each corner, when an adjustment of the mirrors was necessary. A Worthington air-pump, driven by a
Fig. 1.—Ground plan, showing arrangement of mirrors
50 horse-power motor, reduced the pressure in the pipe to about a half-inch in three hours. Most of the measurements were made when the residual air in the pipes had been reduced to a pressure of about half an inch to one inch of mercury. At these pressures the fringes were perfectly steady, and as sharply defined as could be desired.

Fig. 2.—Details of corner box and mirror mounting

The fringes to be measured were those formed by the beams going in opposite directions about the circuit $ADEF$. As a fiducial mark from which to measure the displacement, a second set of fringes was formed by the mirror system $ABCD$. The area inclosed by this circuit was much too small to give a measurable displacement of the fringes, and the shifts actually recorded were those between the central fringes of the two sets.
In general the two sets of fringes will not coincide in position, entirely aside from any question of ether drift or the earth’s rotation, unless the two direct images and the two reflected images of the source are exactly superposed. The central fringes of the set formed by the mirrors of the short circuit will be halfway between the direct and reflected images of the source, and the central fringe of the long line would be halfway between the direct and reflected images if there were no difference due to the earth’s rotation.

To correct for any lack of superposition of the two sets of rays, the observing telescope (a six-inch achromatic objective, and two-inch micrometer eyepiece) was focused on the images of the source (arc or slit) and the apparent displacement of the central fringe of the long circuit, compared with the central fringe of the short circuit, was corrected by an amount equal to the difference in the mean positions of the two images for the two light circuits. The fringes are most conveniently observed in the overlapping cones of light an inch or so inside or outside of the focal plane.

About half of the determinations were made with the arc placed directly in front of the window at A, and about half with a condensing lens, slit, and collimating lens. The second arrangement gave much more light than the first, but there was no apparent difference in the measured displacements.

The calculated value of the displacement on the assumption of a stationary ether as well as in accordance with relativity is

\[ \Delta = \frac{4A \omega \sin \phi}{\lambda V} \]

where \( \Delta \) is the displacement in fringes, \( A \) the area in square kilometers, \( \phi \) the latitude (41°46′), \( V \) the velocity of light, \( \omega \) the angular velocity of the earth, and \( \lambda \) the effective wave-length of the light used. Measurements were made in the laboratory, comparing the fringes produced by the same set of mirrors and the same 20-ampere alternating-current arc, with fringes produced by sodium light from a bit of glass in an oxyhydrogen flame. The light from the arc was reduced to approximately the same intensity as in the experiment at Clearing, by transmitting it through a rather narrow slit in a
rotating disk. The mean of ten determinations gave $\lambda = 5700 \pm 50$ angstroms.

The displacement of the fringes due to the earth’s rotation was measured on many different days, with complete readjustments of the mirrors, with the reflected image sometimes on the right and sometimes on the left of the transmitted image, and by different observers. The deflections were averaged usually in sets of twenty, in the order in which they were taken. The resulting means are given in Table I. The entire set of two hundred and sixty-nine determinations and their distribution about the mean value is shown graphically in Figure 3. The final displacement, expressed as a fraction of a fringe, is

$$0.230 \pm 0.005 \text{ obs.} \quad 0.236 \pm 0.002 \text{ calc.}$$

In view of the difficulty of the observations, this must be taken to mean that the observed and calculated shifts agree within the limits of observational error.

The successful completion of this experiment is due in no small measure to the friendly co-operation of the officials of the City of

<table>
<thead>
<tr>
<th>Observation*</th>
<th>Displacement in Fringes</th>
<th>Number of Observations</th>
<th>Deviation from Mean</th>
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<tbody>
<tr>
<td>1</td>
<td>0.252</td>
<td>20</td>
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<tr>
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<td>5</td>
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<td>20</td>
<td>0.005</td>
</tr>
<tr>
<td>6</td>
<td>0.207</td>
<td>20</td>
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</tr>
<tr>
<td>7</td>
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<td>20</td>
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<tr>
<td>8</td>
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<td>9</td>
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<tr>
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<tr>
<td>12</td>
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<td>20</td>
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<tr>
<td>13</td>
<td>0.230</td>
<td>23</td>
<td>0.000</td>
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Mean 0.230

Total 269

Av. dev. from Mean 0.016

* Observations 1 and 2 corrected for direct image only; 1–6 inclusive, without collimator; 7–13 inclusive, with collimator.
Chicago, Colonel A. A. Sprague, commissioner of public works; Mr. H. L. Lucas, superintendent of the Water Pipe Extension Division, and to Mr. B. W. Cullen, superintendent of the South District of the same Division; to the courtesy of the Clearing

Industrial District, through Mr. Henry H. Porter; to the kindness of the Chicago Telephone Company, and to the hearty and enthusiastic support of the University. We are especially indebted to President Burton, Mr. J. C. Dinsmore, Mr. L. R. Flook, and to Messrs. Thomas O’Donnell, Joseph Purdy, and Charles Stein.

Ryerson Physical Laboratory
University of Chicago
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